

# Lab 7 Worksheet

You are required to hand this worksheet in at the end of the lab.

Name: \_\_\_\_\_ User Name: \_\_\_\_\_

- All the example codes are at `/n/ursa/A288C/alien/python_template`.
- The places you need to put your answer in this worksheet are marked by “\_\_\_\_\_”.

1. **[function]** GRB170827A has a spectrum that can be described by the following equation

$$f(E) = 9.90 \times 10^{-4} \left( \frac{E}{50 \text{ keV}} \right)^{-1.215}, \quad (1)$$

where  $f(E)$  is the photon flux at energy  $E$ , with units of photons  $\text{cm}^{-2} \text{s}^{-1} \text{KeV}^{-1}$ .

Follow example code `function_example.py` to write a function that returns the value of  $f(E)$ . Using your code to find  $f(E)$  at  $E=100 \text{ keV}$ .

$f(100 \text{ keV}) =$  \_\_\_\_\_.

2. **[integration]** Follow the example code `integration.py` and write a code to calculate the the photon flux in the energy range of 15 keV to 150 keV by integrating eq. 1,

$$\text{Photon Flux (15 - 150 keV)} = \int_{15\text{keV}}^{150\text{keV}} 9.90 \times 10^{-4} \left( \frac{E}{50 \text{ keV}} \right)^{-1.215} dE, \quad (2)$$

From your code, photon flux (15-150 keV) = \_\_\_\_\_.

The unit of the photon flux is \_\_\_\_\_.

3. **[sanity check]** Integrate eq. 1 analytically, did you get the same answer as your code? Write down the equations of how you derive the analytical answer:

\_\_\_\_\_

4. **[FITS file and 2d-array]** Follow the example code `fits_file_example.py` and do the following questions:

- (a) Download the 4-channel light curve of GRB170803A from the following link:  
[https://swift.gsfc.nasa.gov/results/batgrbcats/GRB170803A/data\\_product/00766081000-results/lc/sw00766081000b\\_4chan\\_1s.lc](https://swift.gsfc.nasa.gov/results/batgrbcats/GRB170803A/data_product/00766081000-results/lc/sw00766081000b_4chan_1s.lc)
- (b) Print out the time relative to the BAT trigger time  $T_0$ . That is, print out “time -  $T_0$ ”.
- (c) Print the energy bands of the light curve to the screen.

The energy ranges of this light curve are:

\_\_\_\_\_,  
\_\_\_\_\_,  
\_\_\_\_\_,  
\_\_\_\_\_.

- (d) Print out the rate number in 25-50 keV of time between T0 and T0+1 s. (You should find only one value.)  
The rate is \_\_\_\_\_.
5. **[plotting]** Try to run the example code `plot_example.py`, and perform the following changes:
- (a) Change the line style to 'dashed line'.
  - (b) Produce a scatter plot by changing the linestyle to "None", and add the option `marker="o"`.
  - (c) Based on the options in the matplotlib tutorial [https://matplotlib.org/users/plotlib\\_tutorial.html](https://matplotlib.org/users/plotlib_tutorial.html), change the marker to star.
6. **[plot error bars]** Try to run the example code `plot_errorbar.py`, and perform the following changes:
- (a) Produce a scatter plot with error bars with the same set of data.
  - (b) Based on the options in the matplotlib tutorial, change the range in the x-axis to `[-20.0, 30.0]`.
7. **[bar plot (histogram)]** The example code `plot_histogram.py` shows an example of making a bar plot (histogram). Follow this example and use the GRB redshift list [https://swift.gsfc.nasa.gov/results/batgrbcat/summary\\_cflux/summary\\_general\\_info/GRBlist\\_redshift\\_BAT.txt](https://swift.gsfc.nasa.gov/results/batgrbcat/summary_cflux/summary_general_info/GRBlist_redshift_BAT.txt) to make a histogram of redshift distribution with bin size of 0.1.